

Lighting Up The Dark: The Mystery of Dark Matter

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Galaxies



Our galaxy: the Milky Way

~100,000 light years across



Galaxies

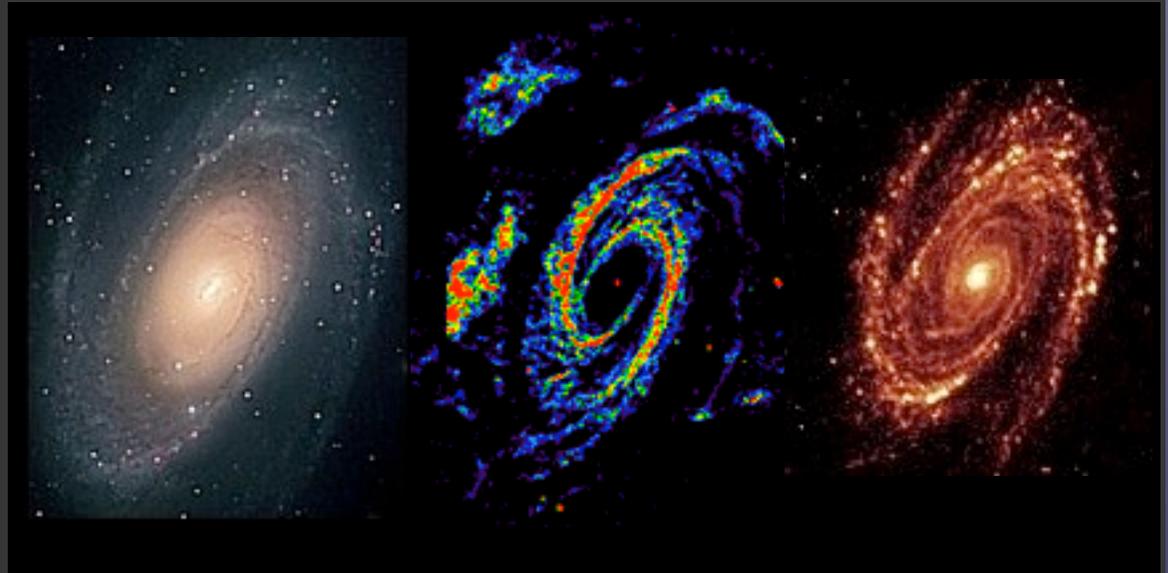


What are galaxies made of?

Stars

Gas

Dust



Optical

Radio

Infrared

What are galaxies made of?

Dark Matter (90%)

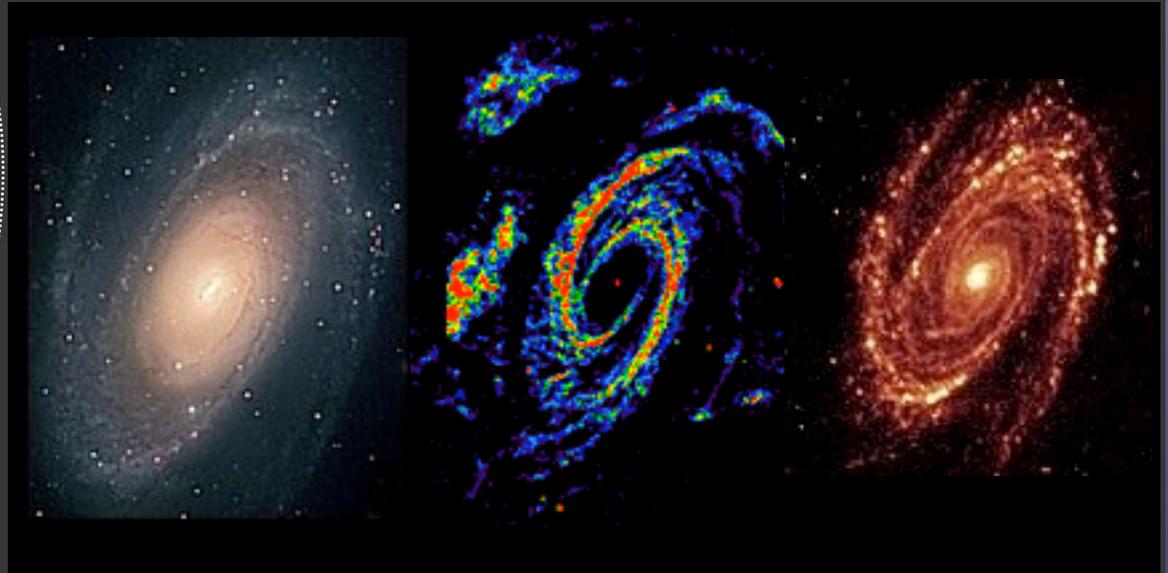
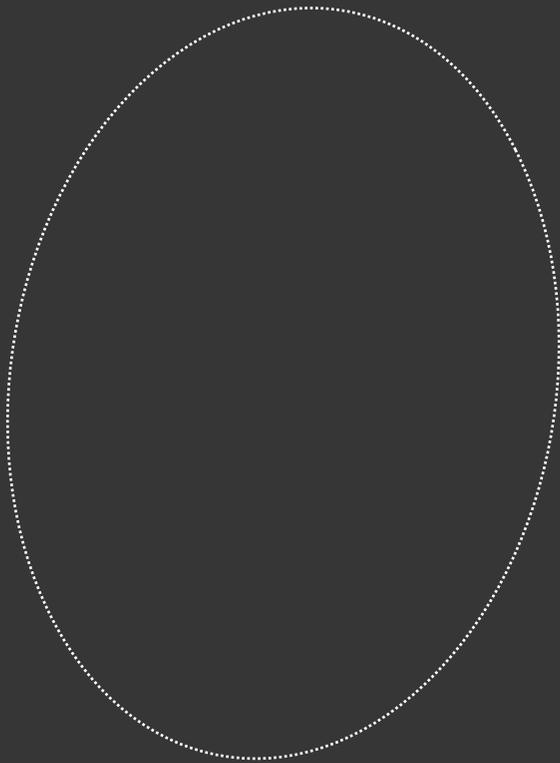
+

Normal Matter (10%)

Stars

Gas

Dust



Optical

Radio

Infrared

Dark Matter



Existence was first postulated in the 1930s from observations of galaxy motions - there was apparently a lot more mass than could be seen!

Fritz Zwicky, Caltech professor

- Dark matter has since been confirmed through observations of:
- fast-moving galaxies in clusters
 - fast rotation of stars and gas in spiral galaxies
 - gravitational lensing (strong and weak)

Dark Matter - What is it?

Unlike normal (baryonic) matter, dark matter doesn't emit or absorb light.

Detectable through its gravitational effects!
It can 'bend' light by bending space.

Likely a slow-moving ('cold'), as yet undiscovered, weakly-interacting massive particle (WIMP).

Not sure exactly what it is, but there are a lot of possible particles it could be. There are on-going experiments in mines underground around the world to detect it directly!

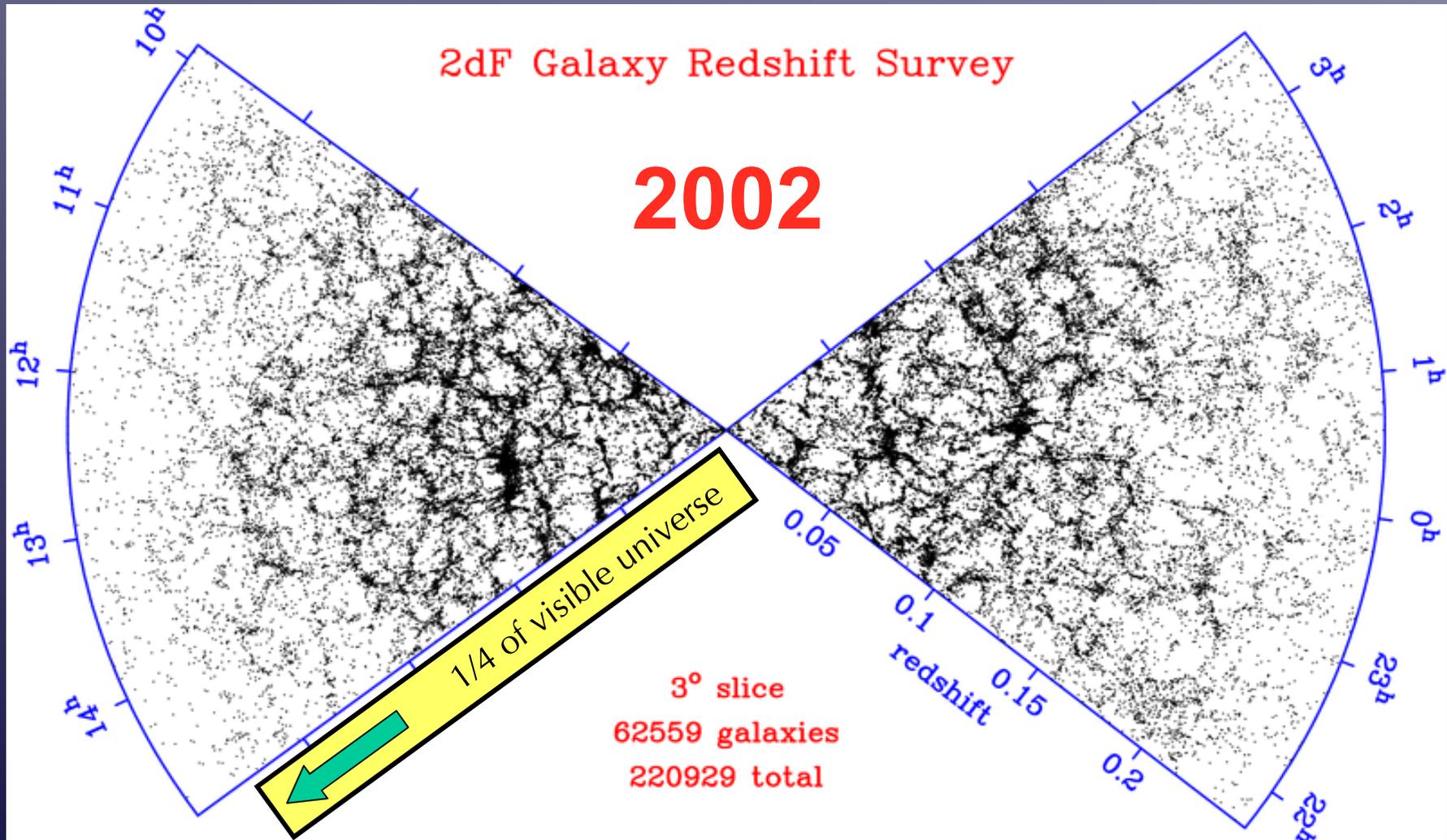
Large-Scale Structure

Galaxies are not evenly distributed in space.

Because of gravity, they tend to form clumps, where many galaxies are found near each other.

This happens on really, really large scales.

Large-Scale Structure



~3 billion light years

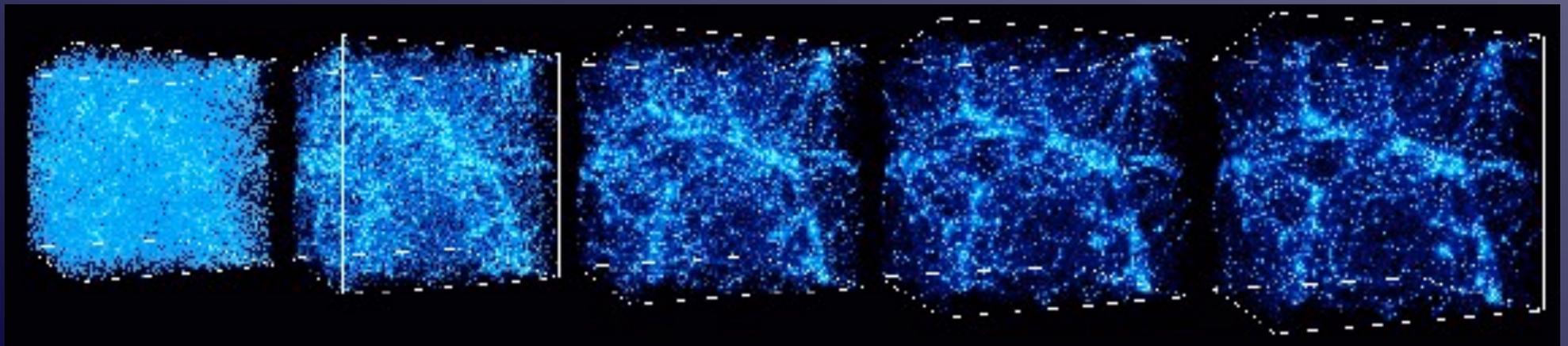
Large-Scale Structure



How did this happen?

2 reasons: quantum mechanics and gravity

- * Very early universe: hot, dense soup of mass/energy
- * Tiny fluctuations: seeds of structure
- * Galaxies formed from initially small amplitude lumps
- * Process of structure formation continues today!



Time \longrightarrow

Dark Energy

Observations in 1998 of type 1a supernovae (exploding stars) and their distances (redshifts) revealed that the Universe is not just expanding but is *accelerating*.



The Nobel prize in physics this year was awarded for this discovery.

Called 'dark energy' as a reference to dark matter.

Counteracts gravity - a mysterious, repulsive force throughout space.



Photo: Roy Kaltschmid, Courtesy: Lawrence Berkeley National Laboratory

Saul Perlmutter



Photo: Selinda Pratten, Australian National University

Brian P. Schmidt



Photo: Homewood Photography

Adam G. Riess

Dark Energy - What is it?

- No one knows!
- Do know that it smoothly fills space, has a constant (or nearly so) density and is repulsive - negative pressure and drives acceleration.
- Could be due to a vacuum energy. Quantum mechanics predicts that empty space isn't empty - lots of short-lived particles coming into and out of existence. These particles provide a net background energy. The more volume of space there is, the more vacuum energy.
- Could be a new force we don't know about yet.
- Our understanding of gravity could be wrong!

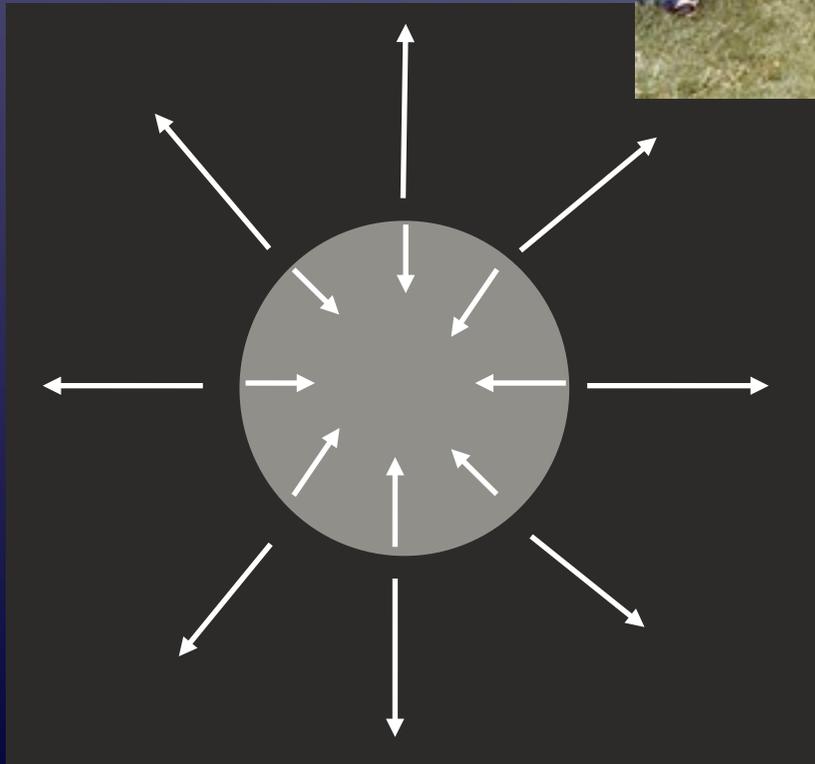
Dark matter and dark energy affect the expansion history of the Universe and therefore affect large-scale structure!

Formation of Structure

Cosmological expansion / dark energy

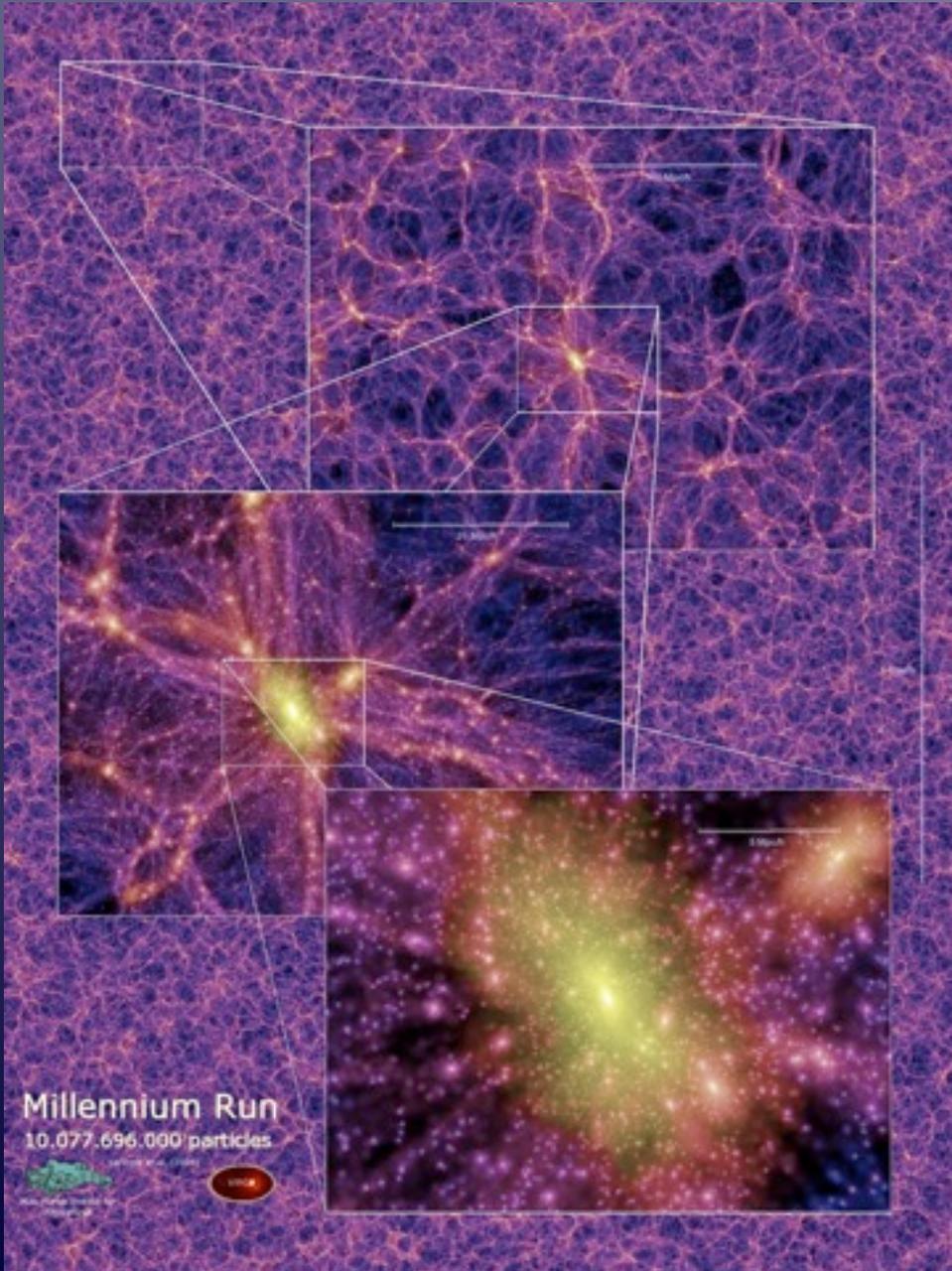


Gravity / dark matter



Overdensities grow into **gravitationally bound** structures. Space between them expands.

Supercomputer Simulation

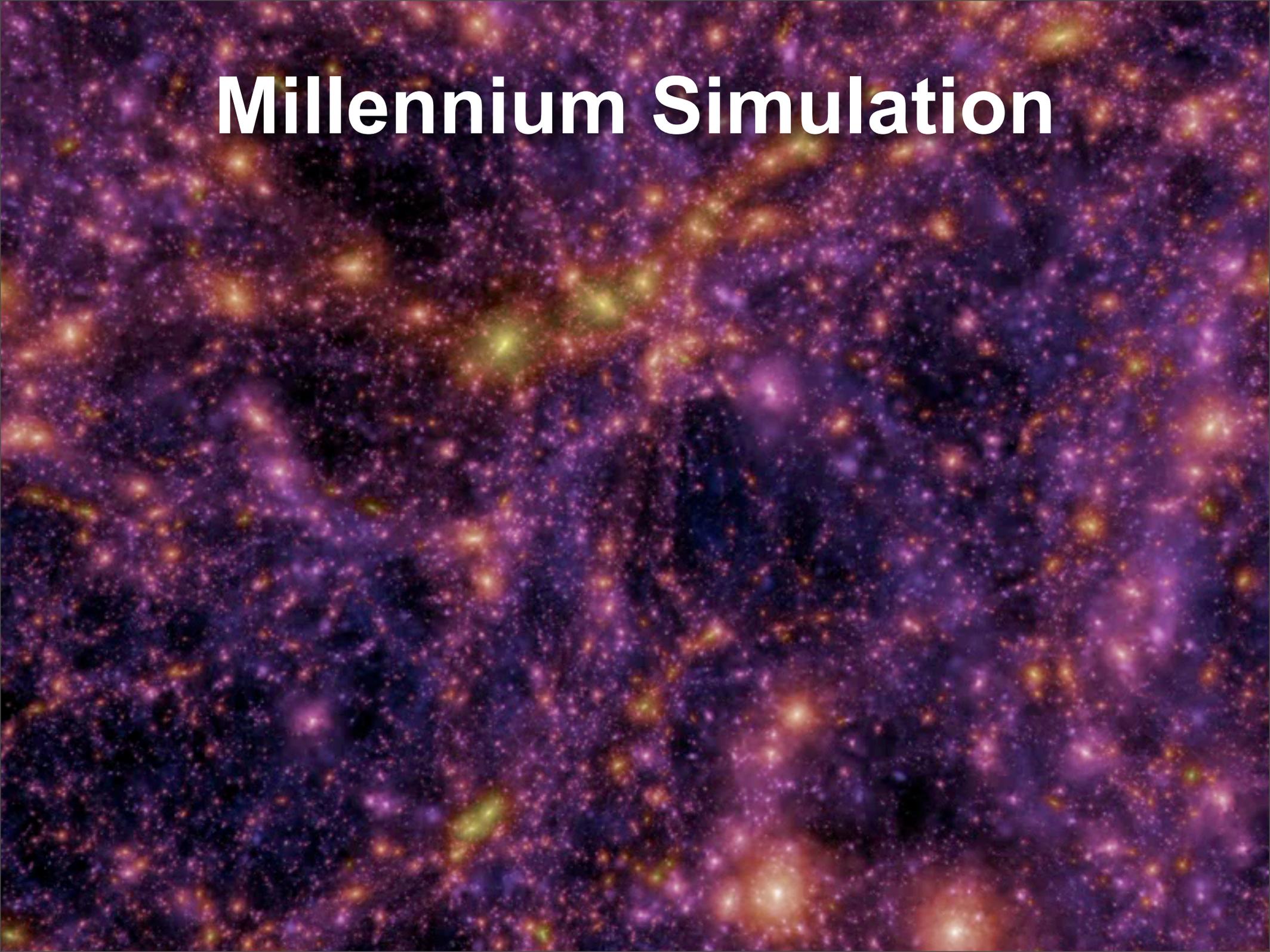


Huge computer simulations of large-scale structure predict the distribution and evolution of dark matter.

Include dark energy and dark matter, but no baryonic/‘normal’ matter - no galaxies!

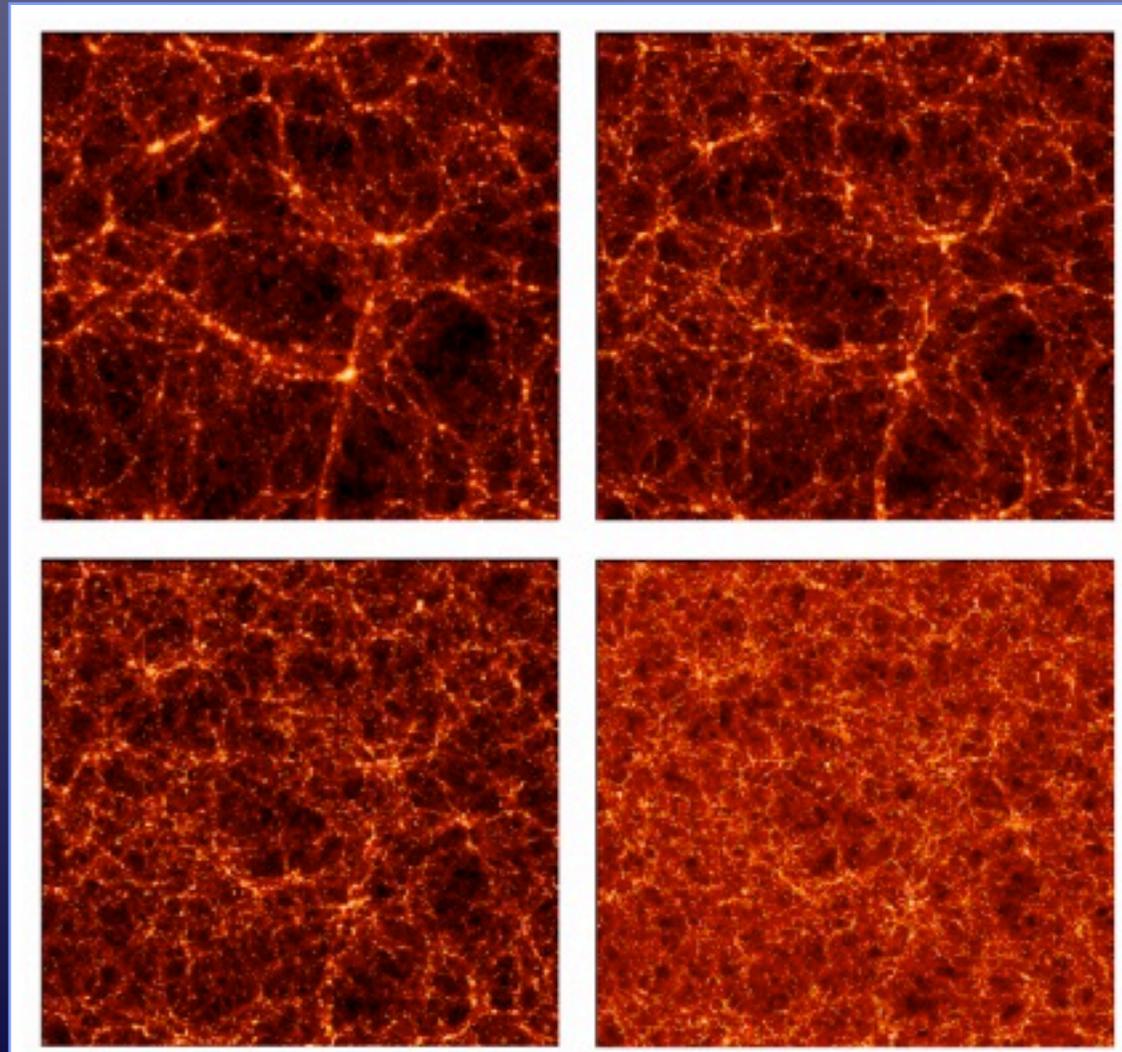
> 10 billion particles
2 billion light years wide
25 Tbytes of stored data

Millennium Simulation



Large-scale Structure Evolution

dark matter evolution:



w/ dark energy
(and dark matter)

w/ a lot of dark matter
(and no dark energy)

Now

7 Gyr ago

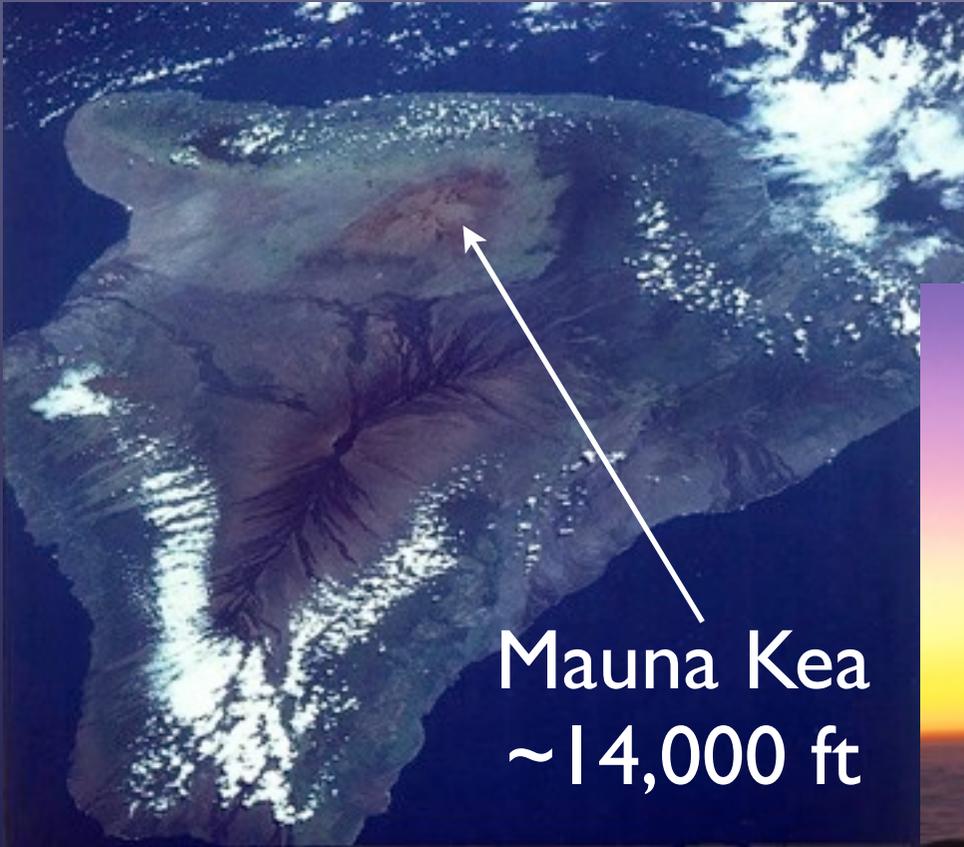
A Survey of Distant Galaxies

Carry out a survey of distant galaxies to test the cosmological model and learn about galaxy evolution.

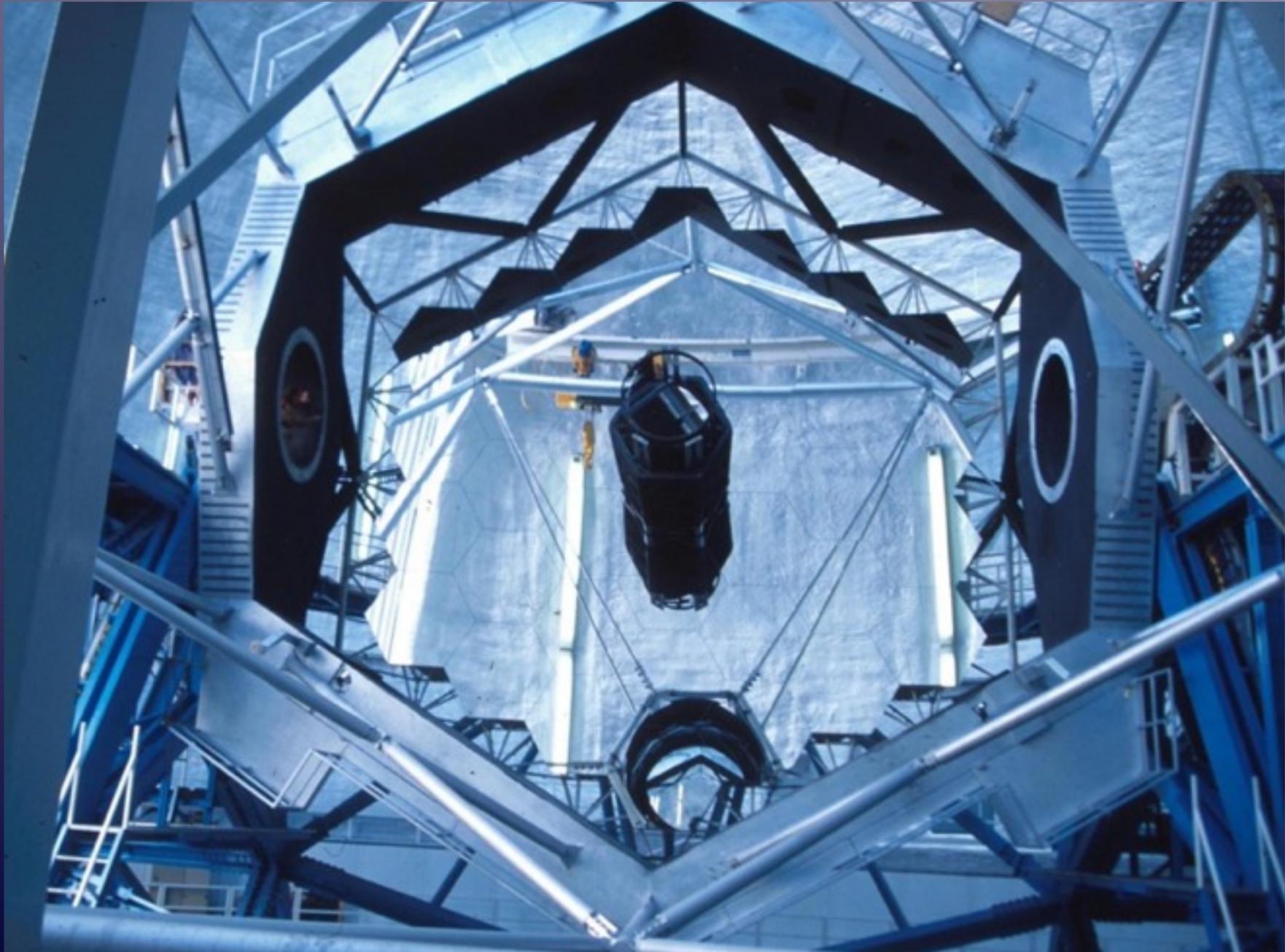
**Survey is called DEEP2 - UC collaboration.
~20 people on the team.**

Use the largest optical telescope in the world to study faint, distant galaxies.

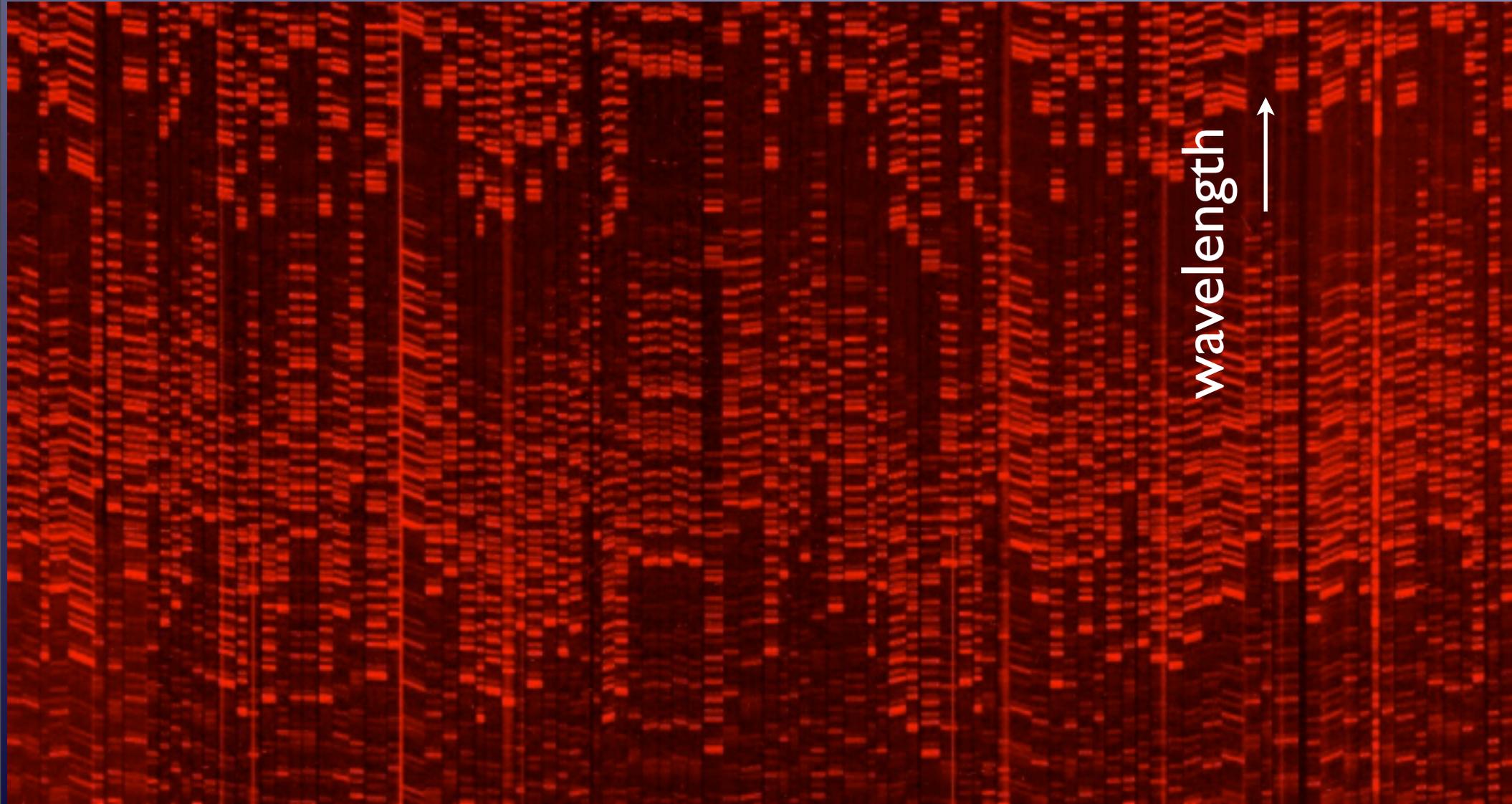
The Power of the Keck Telescope



Keck 10-m telescope

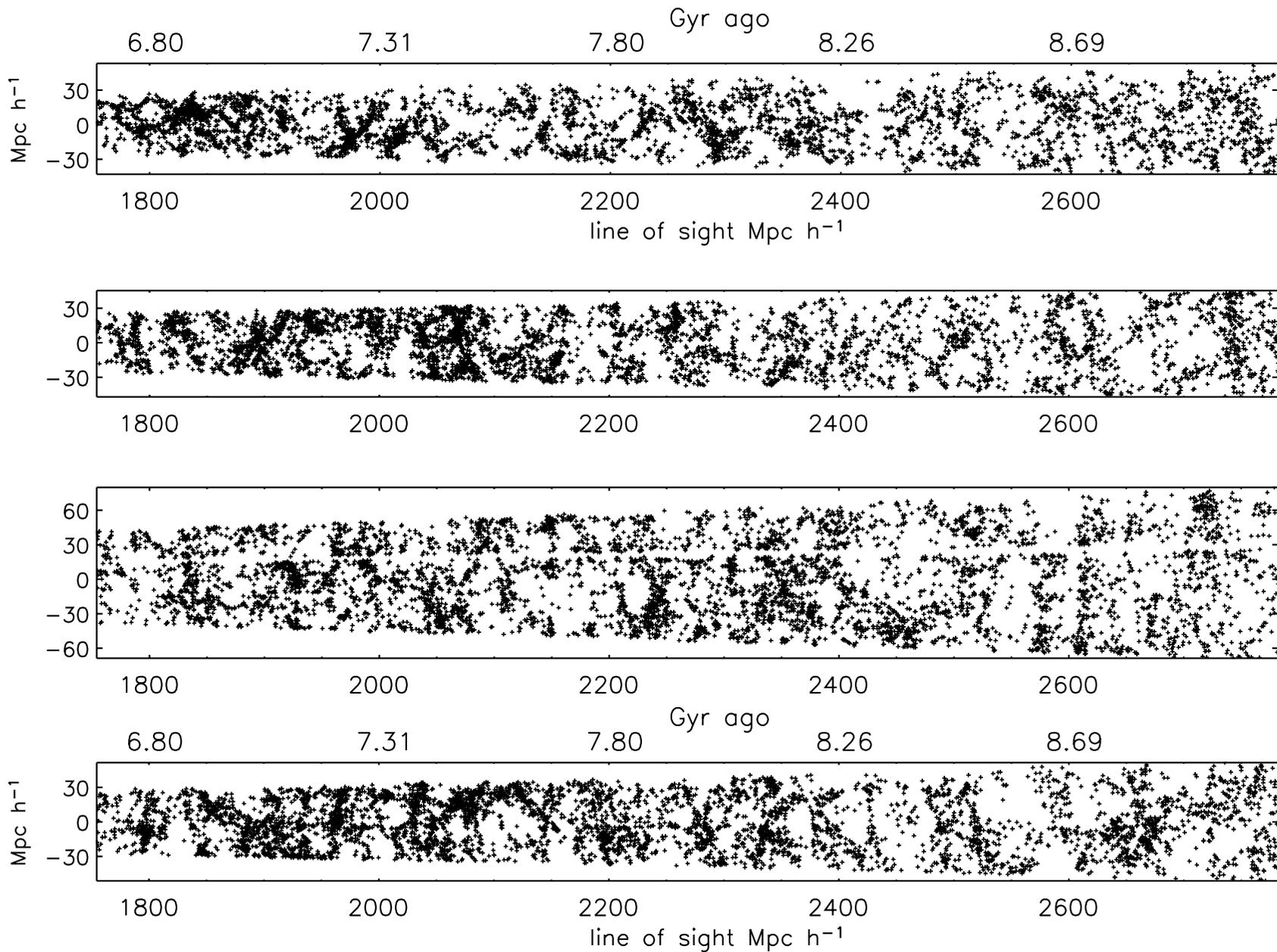


Raw Data - 1/4 of a frame



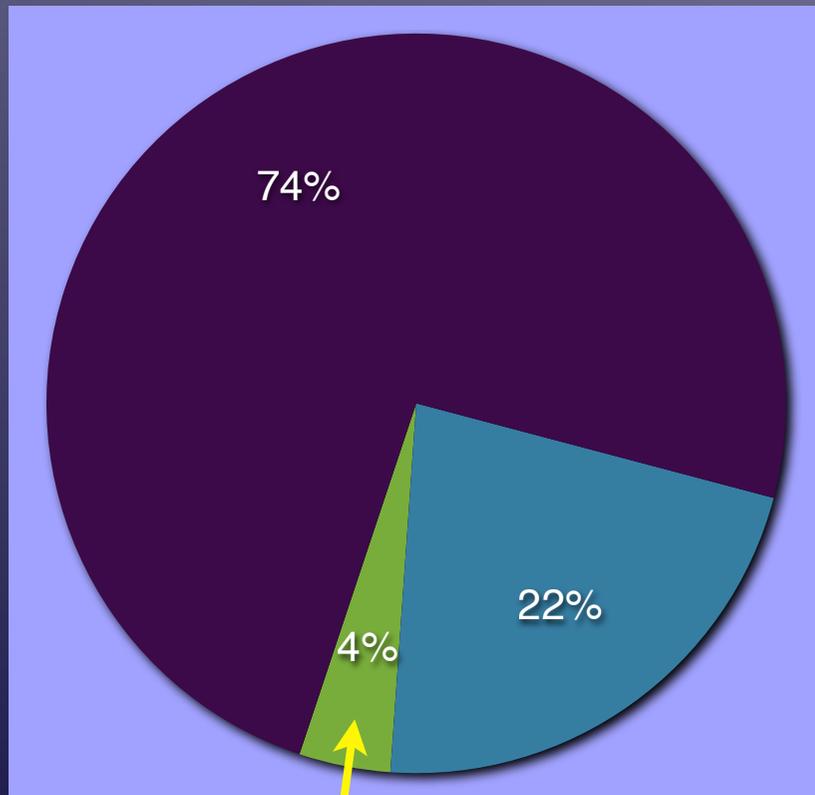
Each column is light from a single galaxy

Distant Galaxy Distribution



How do galaxies cluster?

● Dark Energy ● Dark Matter ● Normal Matter



us!

**Depends on what's in
the Universe:**

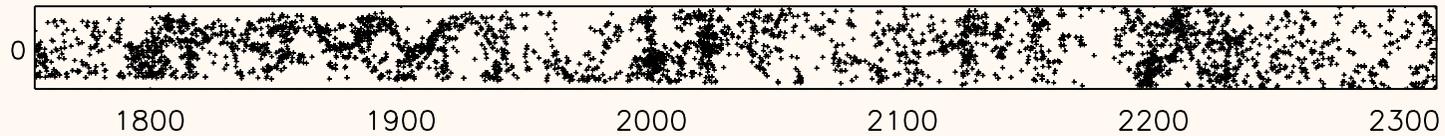
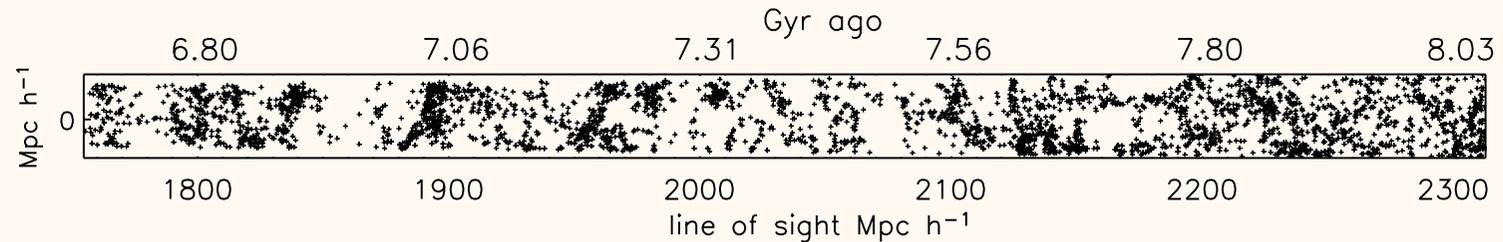
~74% dark energy

~22% dark matter

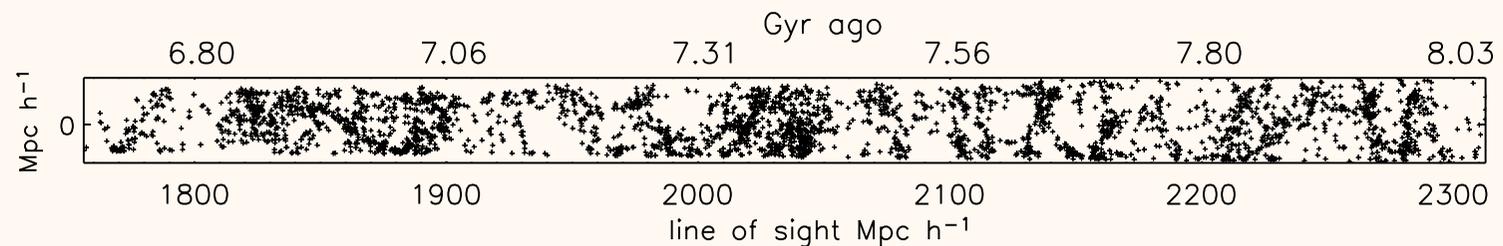
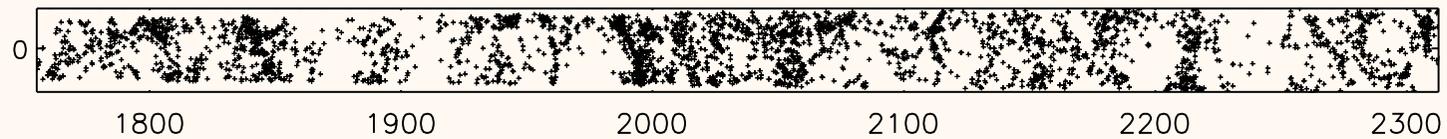
~4% normal matter

Matches Predictions

Agrees w/ the cosmological model!



Computer simulation

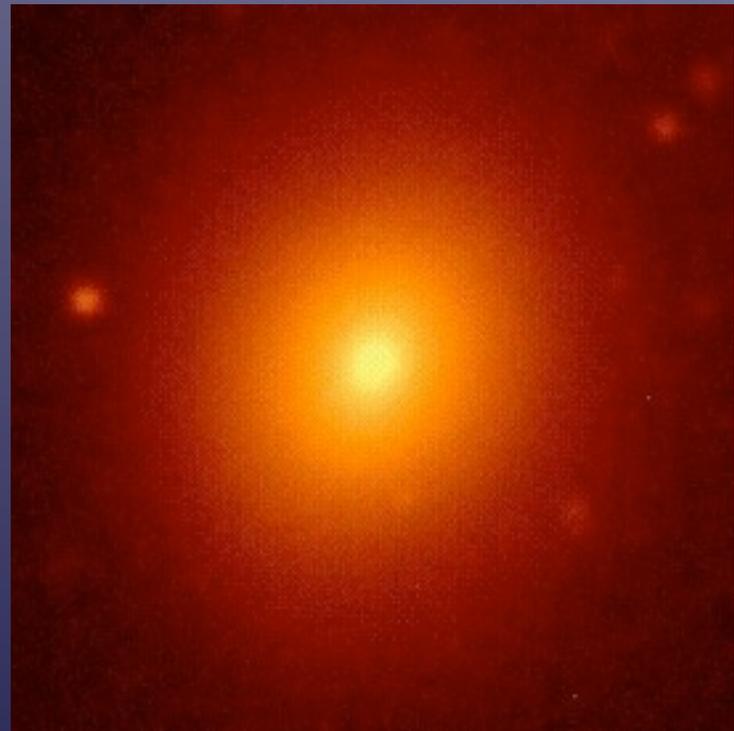


Galaxies Come in 2 Basic Types

spiral, forming stars,
lots of dust and gas,
blue color



elliptical, not forming stars,
little dust and gas,
red color



Add Galaxies

Use 'recipes' for the complicated physics of galaxy formation and evolution

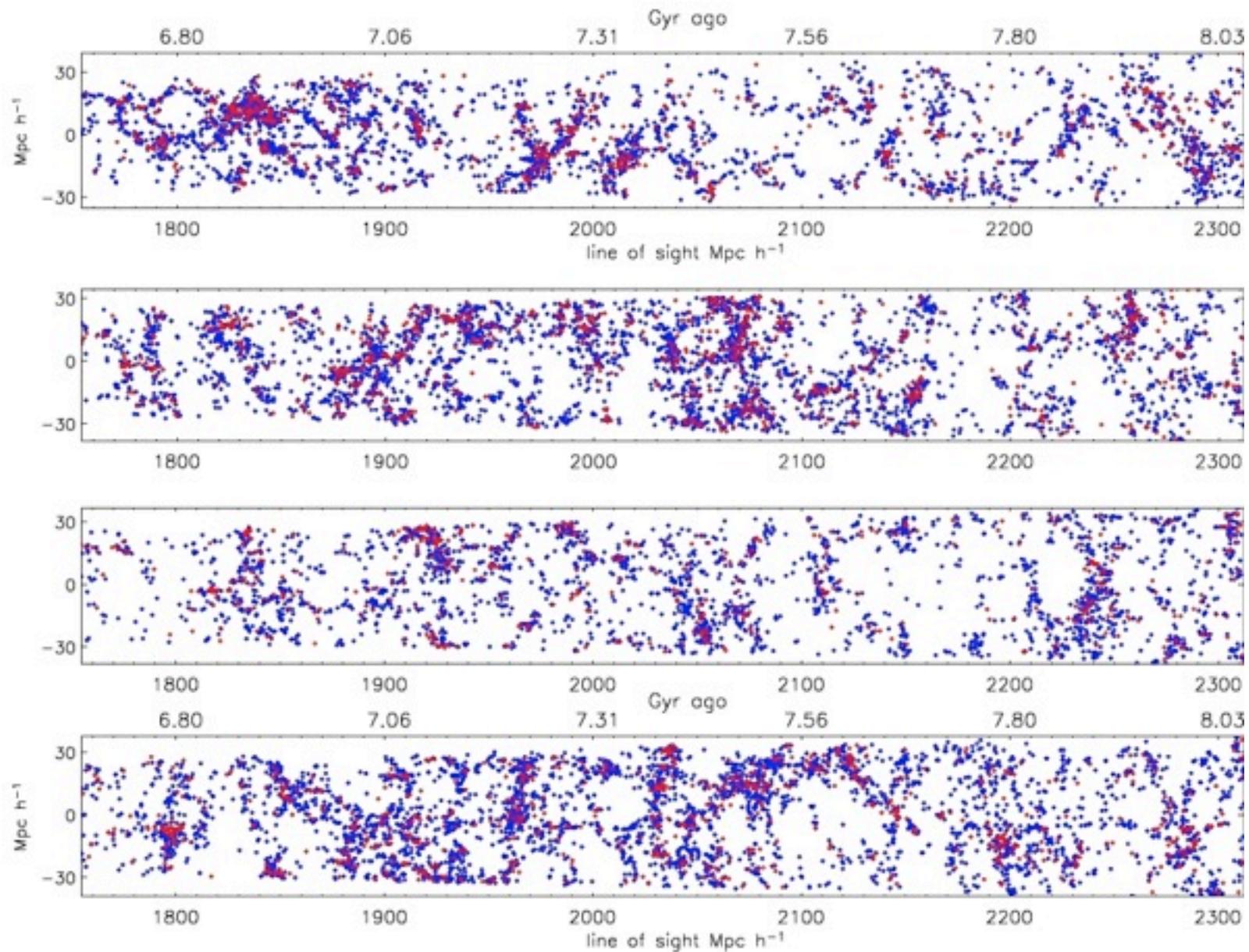


Gas Cooling
Star Formation
Disk creation
Supernovae
Shocks
Outflowing Winds
Black Holes

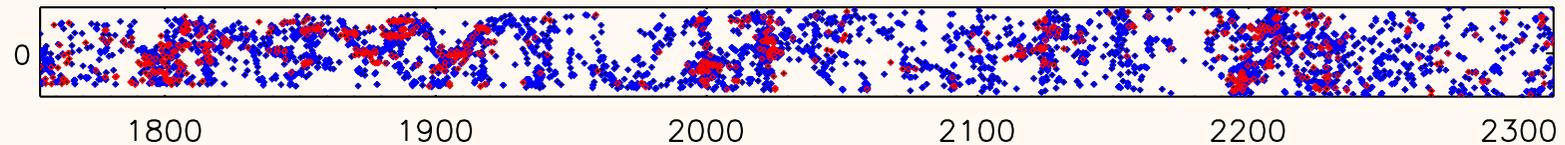
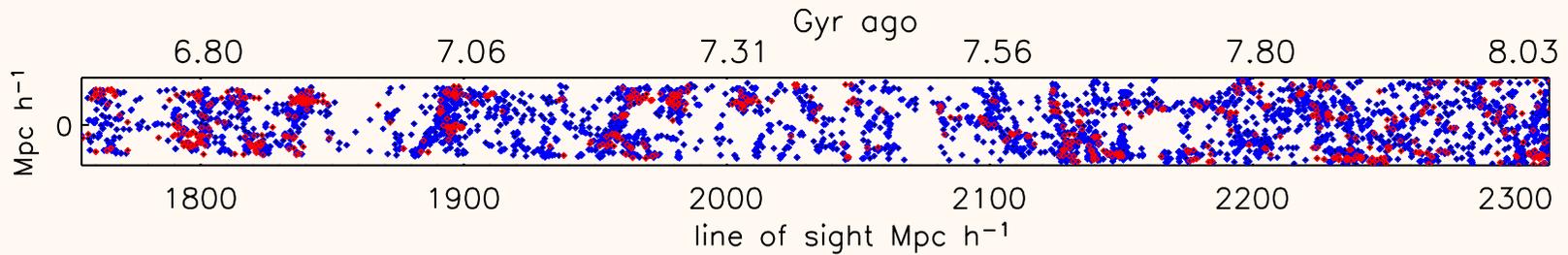
All recipes are tied to observations but have some freedom / wiggle room

Split by Galaxy Type

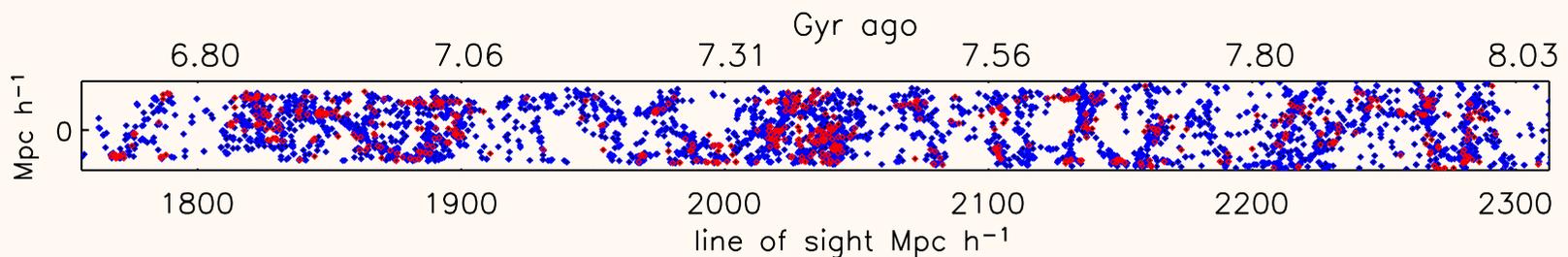
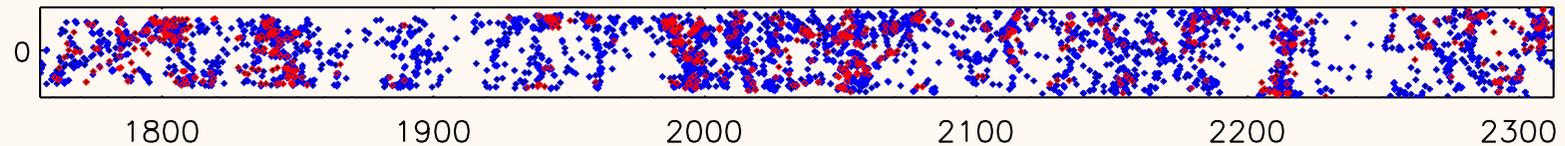
real data:



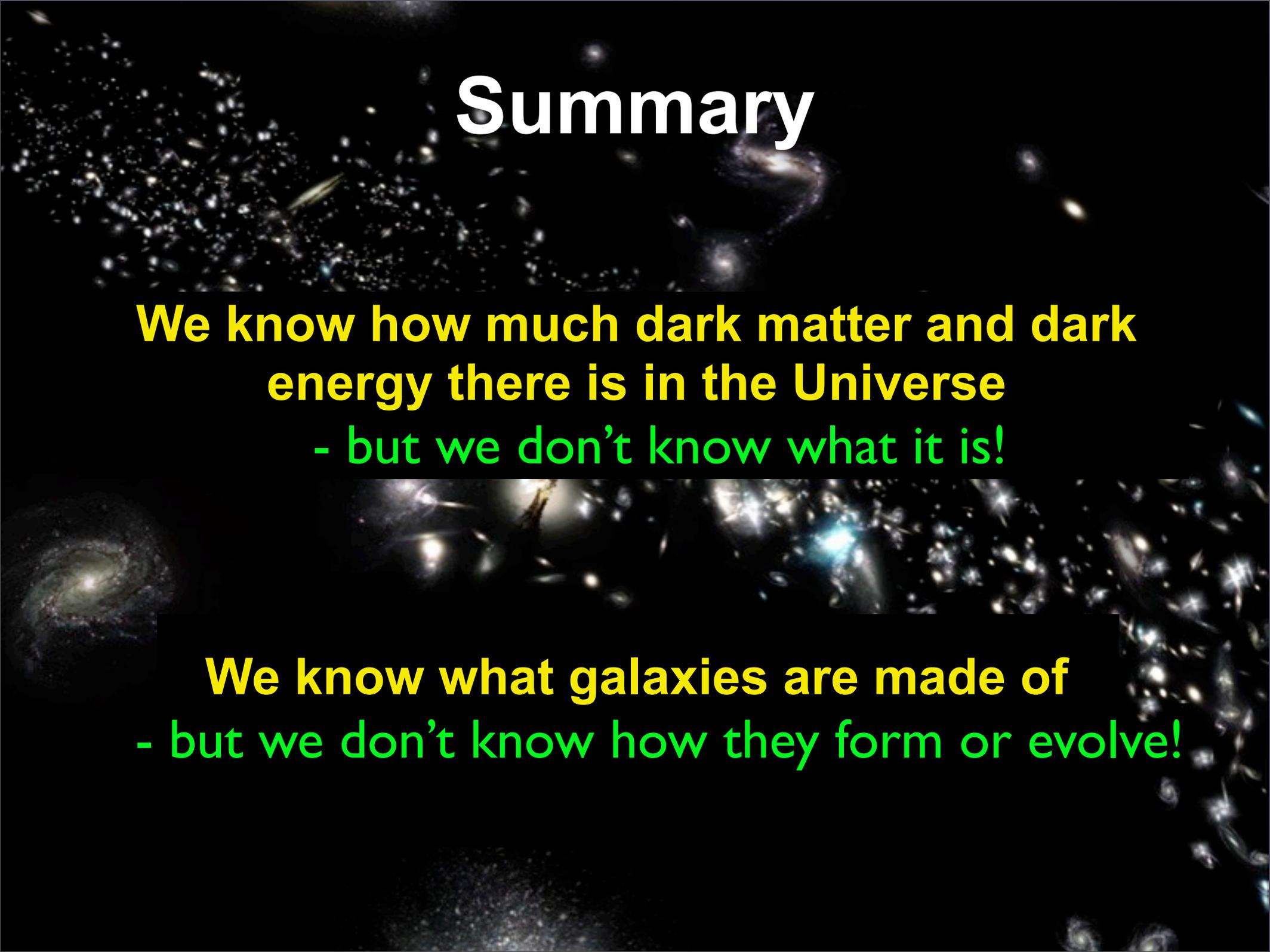
Comparison w/ Simulation



Red galaxies are too clustered,
not enough blue galaxies in dense regions



Summary



We know how much dark matter and dark energy there is in the Universe
- but we don't know what it is!

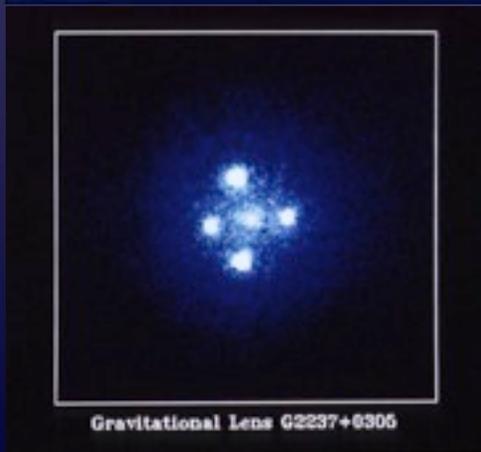
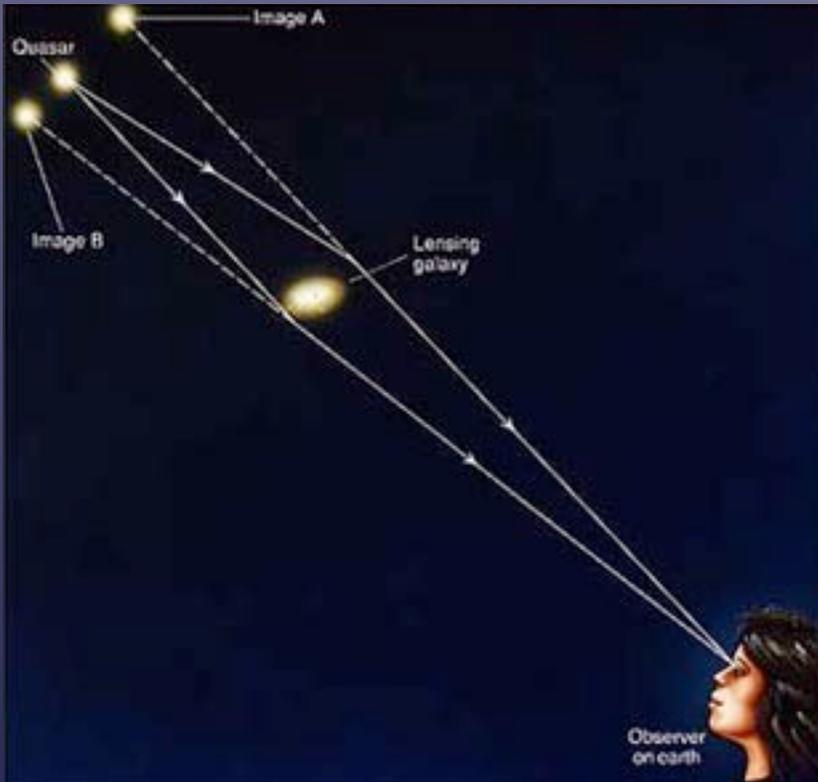
We know what galaxies are made of
- but we don't know how they form or evolve!

Summary

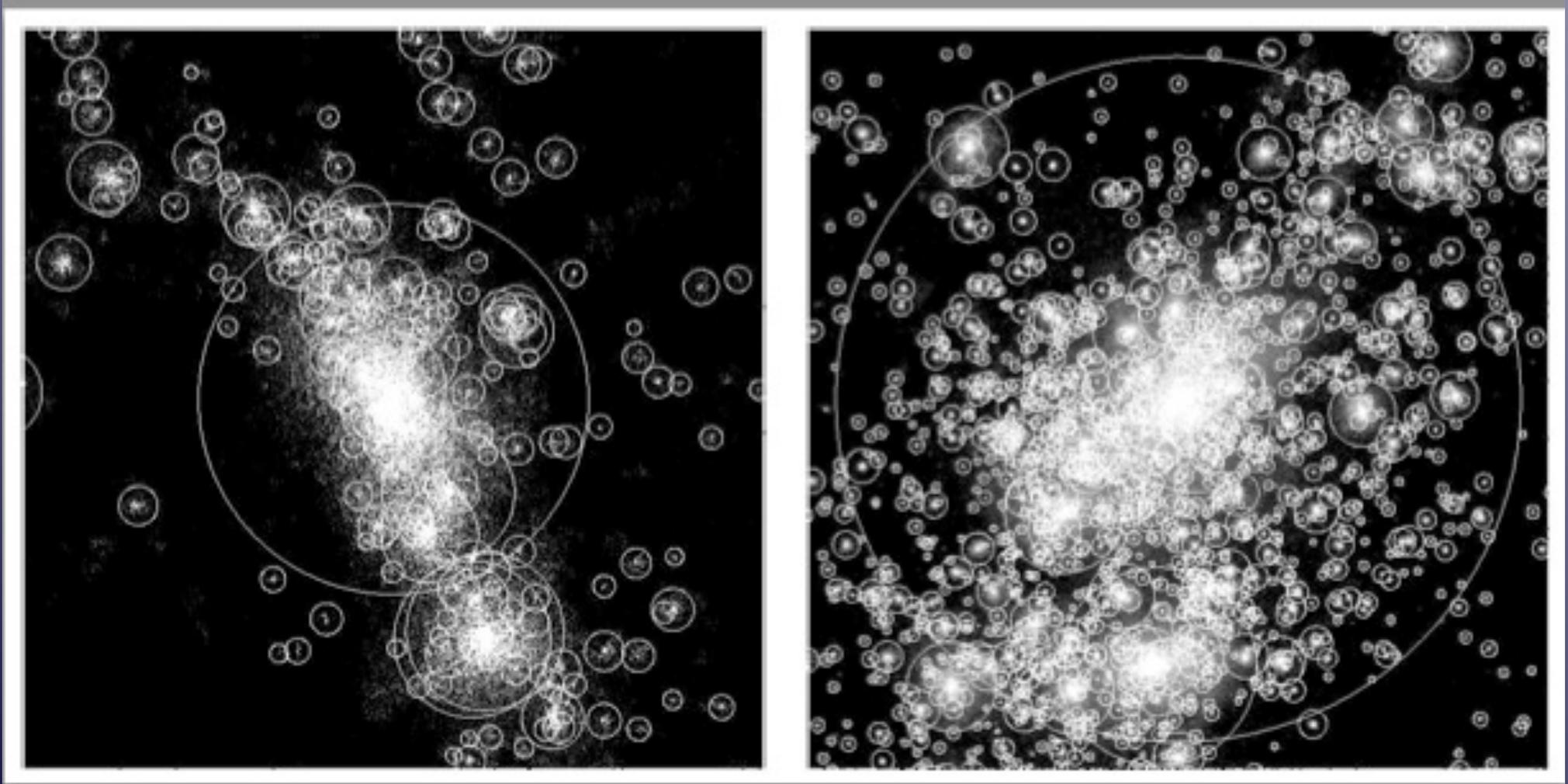
The background of the slide is a composite image of various celestial objects. It features a dense field of small, distant galaxies, some appearing as bright points of light and others as faint, elongated structures. There are also several larger, more prominent galaxies, including a prominent spiral galaxy on the right side and another on the left. The overall scene is set against a deep black background, with the light from the galaxies creating a sense of depth and vastness.

We still have a lot to learn!

Gravitational Lensing



Dark Matter Halos



Overdensities in dark matter distribution collapse through gravity - dark matter halos. Simulations assume one galaxy at the center of each halo and subhalo, above some mass threshold.